Multidisciplinary Workshop on RPA UPMC Paris, 26-29 Jan 2010



How APPROXIMATE IS THE RANDOM PHASE APPROXIMATION? COMPARING RPA AGAINST FULL CONFIGURATION-INTERACTION CALCULATIONS

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I. Stetcu and C. W. Johnson, Phys. Rev. C 66 034301 (2002)

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SYSTEMATIC INVESTIGATION OF RPA OR OUTLINE OF MY TALK!



Review of RPA

Overview of previous "benchmarking" of RPA

Configuration-interaction (CI) calculations and our code SHERPA

Two tests:

(1) Correlation energies

(2) Transitions in RPA and pnRPA

"Collapse" of RPA

Conclusions ... and future work



Review of RPA

RPA models excited states as small oscillations about the mean-field.

One can put RPA in the same framework as configuration-interaction (CI) calculations

using an occupation-space (shell-model) basis

THE NUCLEAR LANDSCAPE





THE NUCLEAR LANDSCAPE



There are many ways to derive the random phase approximation (RPA): eqns of motion, time-dependent Hartree-Fock, linear response...

I prefer quantization of the energy surface:



We often approximate a potential by a harmonic oscillator

In occupation space, the second derivatives are given by the **A** and **B** matrices:

$$A_{mi,nj} \approx \left\langle mi^{-1} \left| H \right| nj^{-1} \right\rangle$$
$$\left| mi^{-1} \right\rangle = \hat{c}_{m}^{+} \hat{c}_{i} \left| HF \right\rangle$$
$$B_{mi,nj} \approx \left\langle mi^{-1} nj^{-1} \left| H \right| HF \right\rangle$$

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For details, see the excellent monograph *The Nuclear Many-Body Problem* by P. Ring and P. Schuck

brought into diagonal form using a Bogoliubov (quasiboson) transformation $E(0) - \frac{1}{2}TrA + \sum_{\lambda} \hbar\Omega_{\lambda} \left(\hat{\beta}_{\lambda}^{+} \hat{\beta}_{\lambda} + \frac{1}{2} \right)$ $\hat{\beta}_{\lambda} = \sum_{i} X_{mi,\lambda} \hat{b}_{mi} - Y_{mi,\lambda} \hat{b}_{mi}^{+}$

and

$$\begin{pmatrix} \mathbf{A} & \mathbf{B} \\ -\mathbf{B}^* & -\mathbf{A}^* \end{pmatrix} \begin{pmatrix} \vec{X}_{\lambda} \\ \vec{Y}_{\lambda} \end{pmatrix} = \hbar \Omega_{\lambda} \begin{pmatrix} \vec{X}_{\lambda} \\ \vec{Y}_{\lambda} \end{pmatrix}$$

The RPA matrix equation

There is a correction to the Hartree-Fock energy due to "zero-point motion" or, correlations among the nucleons:

$$E_{RPA} = E(0) - \frac{1}{2}Tr\mathbf{A} + \sum_{\lambda} \hbar\Omega_{\lambda}\left(\frac{1}{2}\right)$$



Historical validation of RPA



Despite its widespread use, RPA has generally been only tested against toy models (in nuclear physics)

A typical exampe are Lipkin-type models:

Parikh & Rowe, Phys Rev **175** (1968) 1293 Hagino & Bertsch, PRC C **61** (2000) 024307

acts like parity conservation



N particles, each either up or down... simple quasispin Hamiltonian The 2-body interaction promotes or demotes 2 particles at a time

Your basic Lipkin model has only 2 independent parameters: N, the number of particles and ratio of 2-body to single-particle spitting, V/ ϵ



Despite its widespread use, RPA has generally been only tested against toy models

A typical exampe are Lipkin-type models:

$$\hat{H}_{LMG} = \varepsilon \hat{J}_z - V \left(\hat{J}_+^2 + \hat{J}_-^2 \right) \begin{bmatrix} 175 \ (1968) \ 1293 \\ 61 \ (2000) \ 024307 \\ acts \ like \ parity \ conservation \end{bmatrix}$$

$$\hat{J}_{z} = \frac{1}{2} \sum_{i} \hat{a}_{i\uparrow}^{+} \hat{a}_{i\uparrow} - \hat{a}_{i\downarrow}^{+} \hat{a}_{i\downarrow}$$
 raction promotes or les at a time set of the set of





I'll return to this point later...



Other tests of RPA...

two-level pairing	Hagino & Bertsch, Nucl. Phys. A679 (2000) 163
schematic interaction in small SM space	Ullah & Rowe, Phys. Rev. 188 (1969) 1640.
+ handful of others	

+ nanotul of others...



The Shell-Model Basis and configuration-interaction calculations and a flexible RPA code in the shell model

Diagonalization of a Hamiltonian in a shell-model basis (a.k.a. configuration-interaction or CI) yields **"exact"** (for that space) and **nontrivial** numerical results

Let's compare RPA against these numerical CI results

How a CI code works



Nuclear CI codes, such as OXBASH, ANTOINE, or REDSTICK/BIGSTICK, (also called 'shell-model codes' in the nuclear structure community) writes the Schrodinger eqn as a matrix eigenvalue equations

One defines a single-particle basis ...



...selects a valence space...

....puts in valence nucleons...

...possibly assuming an inert core.

How a CI code works



The basis is the set of Slater determinants of all (sort of) possible configurations in the valence space



The interaction Hamiltonian is specified as *single-particle energies* plus *two-body matrix elements* (the "residual interaction")

These are read in to the program as a list of numbers

How a Clode works We solve $\mathbf{H}|\Psi\rangle = E|\Psi\rangle$

Computational Nuclear Structure

in a large but finite-dimension space (eigenvalue problem)

i = $(0s_{1/2,1/2})_{\pi}$, etc.

How A Cl code works $\hat{H} = \sum_{i} \varepsilon_{i} \hat{a}_{i}^{\dagger} \hat{a}_{i} + \frac{1}{4} \sum_{ij \, kl} V_{ij \, kl} \hat{a}_{i}^{\dagger} \hat{a}_{j}^{\dagger} \hat{a}_{l} \hat{a}_{j}$ $V_{ij \, kl} = \int dx dx' \phi_{i}^{*}(x) \phi_{j}^{*}(x') V(x, x') \left[\phi_{k}(x) \phi_{l}(x') - \phi_{l}(x) \phi_{k}(x') \right]$

The V_{ijkl} enter the code as *pre-computed numbers*, so there is *no* limitation on the form of the interaction.

How a CI code works



The hard part is actually computing *efficiently* the many-body matrix elements from the two-body matrix elements

That is,
$$ig\langle lpha ig| \hat{H} ig| eta ig
angle$$
 from the V $_{
m ijkl}$

The final result is the low-lying energy spectrum and the corresponding wavefunctions (the coefficients in the Slater determinant basis)



Diagonalization of a Hamiltonian in a shell-model basis yields **"exact"** (for that space) and **nontrivial** numerical results

I have an idea! Let's write an RPA code using exactly the same shell-model input!



SHEII-model RPA code (Stetcu PhD LSU 2003)

Shell-model input compatible with REDSTICK: list of single-particle orbits (Os_{1/2}, Op_{3/2} etc.) list of two-body matrix elements < ab; JT |H|cd;JT > fair to compare output with REDSTICK results

Fully self-consistent Hartree-Fock:

no restrictions on Slater determinant \rightarrow arbitrary deformations within model space (except, wfns purely real)

Standard RPA:

solve matrix RPA equations see rotation of deformed HF state as zero-frequency modes; option to do pnRPA



RPA Correlation Energies

(g.s. Binding energies beyond the mean-field)

RESULTS: CORRELATION ENERGIES



Poorest results for single-species calculations (oxygen, calcium isotopes)

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uclear Structure



RESULTS: CORRELATION ENERGIES

space # nuclides	rms err (keV)
sd (p+n) 41	870
sd oxygen 6	1800
pf (p+n) 11	480
pf calcium 7	730



Transitions





We also looked at Gamow-Teller transitions in pnRPA

here there have been previous detailed comparisons with the shell model, but using spherical pnQRPA

the central question: which is more important,

pairing or deformation?

What is pnRPA?



So far we had separate proton and neutron Slater determinants

$$|\Psi\rangle = |\psi\rangle_{\pi} |\psi\rangle_{\nu}$$

The particle-hole operators conserved charge:

$$X_{\lambda,mi}^{\pi}\pi_{m}^{+}\pi_{i}-Y_{\lambda,mi}^{\pi}\pi_{i}^{+}\pi_{m}$$

pn operators change charge

$$X_{\lambda,mi}^{pn}\pi_m^+\nu_i - Y_{\lambda,mi}^{pn}\pi_i^+\nu_m$$

OTHER'S PREVIOUS WORK: PN-QRPA



A number of papers compared **spherical** pn-QRPA against "exact" shell –model calculations of Gamow-Teller strengths







FIG. 1. Running sum of the Gamow-Teller strength $B(GT^+, E_m)$ for ⁴⁸Ti \rightarrow ⁴⁸Sc as a function of the ⁴⁶Sc 1⁺ excitation energies relative to the ⁴⁶Ti ground state.

Laurtizen, Nucl Phys A489 (1988) 237.

Zhao & Brown, PRC 47 (1993) 2641

Running sum of GT strength

OTHER'S PREVIOUS WORK: PN-QRPA



Most likely explanation: pn-QRPA fails to sufficiently smear the Fermi surface insufficient fragmentation of GT strength





Auerbach, Bertsch, Brown & Zhao, Nucl Phys A556 (1993) 190

OUR CALCULATIONS: DEFORMED PN-RPA



We redid this work, eschewing pairing correlations in favor of unrestricted deformations



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Not only deformation, but triaxiality improves the result



"Collapse" of RPA

at "phase transitions"



Example from the Lipkin model....



Example f





Do we see this with SHERPA?



I know what's happening! I wrote about it in D.Thouless, Nucl. Phys. **22**, 78 (1961)



No collapse of RPA! What's going on?



There are *first-order* and *second-order* transitions!







Quadrupole shape transitions are first order

Lipkin model is 2nd order and is more analogous to mixing of parity across major shells

Example: 0p1/2-0d5/2 model space displays true "collapse"





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CONCLUSIONS

We have realized RPA in a non-trivial shell model framework

Tests of RPA show it to be a modest approximation to the full many-body diagonalization

We have also investigated "collapse" of RPA

We are extending this work to generator coordinate calculations, HFB+QRPA (especially for neutrinoless double-beta decay), and possibly extensions of RPA, e.g. second RPA, etc.





Let *F* be a transition operator; then the <u>energy-weighted sum rule</u> states that

$$\frac{1}{2}\langle HF|\hat{F}\hat{H},\hat{F}HF\rangle = \sum_{\lambda}\hbar\Omega_{\lambda}\left|\langle 0|\hat{F}|\lambda\rangle_{RPA}\right|^{2}$$

This theorem is proven in many text-books...but is wrong!





The "proof" assumes no Goldstone (zero-energy) modes

if one rederives it using those Goldstone modes one gets a correction

$$\frac{1}{2} \langle HF | \left[\hat{F} \left[\hat{H}, \hat{F} \right] HF \rangle = \sum_{\lambda} \hbar \Omega_{\lambda} | \langle 0 | \hat{F} | \lambda \rangle_{RPA} |^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} + \sum_{\nu, \Omega=0} \frac{1}{2M_{\nu}} | \left[F, \hat{P}_{\nu} \right]^{2} +$$

 $\frac{1}{2}\langle HH$

+ $\sum_{\nu \in \Omega} \frac{1}{2M} \left[F, \hat{P}_{\nu} \right]^{2}$

The missing strength can be interpreted as transitions within a rotational band (that is, within the *intrinsic state*) while RPA models transitions within a vibrational band

This is bolstered by the fact that we see missing strength (in even-even nuclides) for E2 transitions but not for, say, spin-flip (ΔJ=1) transitions (because rotational band only allows ΔJ ≥2)

<u>anives it using those</u>

Correction term (Stetcu, 2003) Computational

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The missing strength can be interpreted as transitions within a 5

within the *intrinsic* transitions with

In addition, if we choose a spherical state (no Goldstone modes) rather than a deformed state, we regain the missing strength



Computational

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